AHEAD-2020 WP14 Tasks 14.6 & 14.8

- Blind-Line-Search and ID (14.6)
- Time-Evolving Photoionization (14.8)

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Blind Line Search Routine

- Fortran-90 Routine working in either HEASOFT or CIAO environments
- Extremely fast on 30,000 channel spectra (e.g. Athena-XIFU)
- Reads in:
 - RMF/RSP (FITS format)
 - Continuum-Normalized data in sigma (QDP|ASCII: e.g. from XSPEC or Sherpa)
 - Output Filenames prefix
 - (line statistical significance threshold in sigma + verbosity: default σ =4, v=0)
- Extracts LSF, computes LSF FWHM (i.e. spectral resolution)
- Scans the data searching for positive/negative unresolved, quasi-symmetric, line-like features, with integrated significance
 exceeding the threshold
- Reads in atomic-database (currently Verner et al., 2006) and looks for possible line-IDs (WiP)
- Writes out:

Number of Emission Lines:

- File containing list of 'detected' lines' energy, peak-significance, integrated-significance
- File containing possible redshift IDs (WiP)

nicastro@<u>Falcao.</u>local% bls_v2 athena_xifu_1469_onaxis_pitch249um_v20160401.rsp LSF-convolved_Delchi_28_1.5e7.dat 4 0

| Response Filename: athena_xifu_1469_onaxi Continuum-Normalized Filename: LSF-convol Output Filename: 4 Line Threshold Significance: 0 | ls_pitch249um_v20160401.rsp lved_Delchi_28_1.5e7.dat |
|--|---|
| Using default Verbosity: O | |
| Number of Absorption Lines: 18 | 3 |

2

BLS Example: "Observation"

Line of Sight Extracted from Cen & Ostriker (2006) Hydro-Dynamical Simulations

Column Density

Redshift

Metallicity

| nicastro@Falcao | .local% more | 10S_28_WOZ | _WHIM.OUT | | | | | | |
|-----------------|---------------|------------|---------------|---------------------|---------------|------------------|-------------------|-----------------------|------------------|
| No. of Filamen | t, Last row o | f the Grou | p in Original | File; No of rows in | Group; T(in K | (); NH(in cm-2); | Turbolence (in km | ı∕s); Redshift; log(U |); Density (at t |
| he given z); Me | tallicity (co | mpared to | Solar) | | | | | | |
| 1 | 3847 | 1 | 1379361.25 | 3.36197258E+18 | 261.441772 | 0.109475948 | -1.93419802 | 1.08900003E-05 | 9.50078666E-02 |
| 2 | 4908 | 1 | 729345.250 | 3.23784322E+18 | 190.108902 | 0.140729979 | -1.91785955 | 1.04900000E-05 | 0.201572582 |
| 3 | 4915 | 3 | 688116.562 | 3.43952801E+18 | 184.602371 | 0.140762627 | -1.94454646 | 1.11622676E-05 | 0.506648600 |
| 4 | 4919 | 2 | 907217.188 | 3.20123498E+18 | 212.001709 | 0.140754178 | -1.92046666 | 1.05629624E-05 | 0.220032334 |
| 5 | 9167 | 3 | 206526.281 | 3.33116976E+18 | 101.091904 | 0.273898929 | -1.94133806 | 1.10789997E-05 | 1.66198611E-02 |
| 6 | 13305 | 6 | 214686.875 | 1.11909778E+19 | 103.111465 | 0.416768640 | -2.30846786 | 3.37445308E-05 | 2.01087464E-02 |
| 7 | 13919 | 25 | 5417012.50 | 1.18929300E+20 | 515.280945 | 0.436983317 | -2.99046493 | 7.81381386E-04 | 0.115085848 |
| 8 | 15462 | 1 | 115829.562 | 1.46608968E+19 | 75.7609711 | 0.495346159 | -2.57376432 | 4.75099987E-05 | 1.01446602E-02 |
| 9 | 17371 | 7 | 451797.094 | 5.50937954E+18 | 149.567551 | 0.569676399 | -2.06023479 | 1.45867716E-05 | 0.191525415 |
| 10 | 17940 | 4 | 143990.109 | 4.48879305E+18 | 84.4650497 | 0.590543449 | -2.10365844 | 1.60944182E-05 | 1.46267880E-02 |
| 11 | 18372 | 4 | 398528.188 | 4.00824627E+18 | 140.527100 | 0.608431876 | -2.02583265 | 1.34562370E-05 | 2.12974884E-02 |
| 12 | 18414 | 6 | 286541.844 | 1.97825693E+19 | 119.021568 | 0.609701395 | -1.99990737 | 1.45019585E-05 | 0.195985019 |
| 13 | 18419 | 4 | 371265.344 | 1.66751582E+19 | 135.493530 | 0.610535920 | -2.91440797 | 1.06934771E-04 | 0.309740871 |
| 14 | 19595 | 3 | 1126450.88 | 3.17020483E+18 | 236.260986 | 0.659921110 | -1.91196179 | 1.03496823E-05 | 0.154758915 |
| 15 | 21110 | 5 | 148218.734 | 3.50232332E+18 | 85.6894989 | 0.724813461 | -1.91181695 | 1.03493585E-05 | 7.48509616E-02 |





BLS Example: output

Detected Line-list

| # Number of Ab | sorption Lines: | 17 | |
|------------------|-------------------|---------------|------------|
| # Number of En | nission Lines: | 2 | |
| # E (keV), Sigma | -Peak, Sigma, Sig | gma–Threshold | |
| # Negative Line | S: | | |
| 0.267800003 | 1.59371269 | 4.03757238 | 4.00000000 |
| 0.356200010 | 7.09524679 | 17.2754307 | 4.00000000 |
| 0.413399994 | 1.77576268 | 4.03026819 | 4.00000000 |
| 0.454200000 | 10.0198488 | 24.6709194 | 4.00000000 |
| 0.503000021 | 8.24497032 | 20.4788818 | 4.00000000 |
| 0.564999998 | 2.81150985 | 7.20422220 | 4.00000000 |
| 0.573800027 | 15.4453526 | 39.6384010 | 4.00000000 |
| 0.606999993 | 3.29755354 | 8.53451443 | 4.00000000 |
| 0.623799980 | 1.95347917 | 4.41314697 | 4.00000000 |
| 0.640600026 | 5.42002916 | 7.90693521 | 4.00000000 |
| 0.641399980 | 5.75636578 | 14.7684107 | 4.00000000 |
| 0.710600019 | 7.27873373 | 18.1650715 | 4.00000000 |
| 0.711399972 | 6.76571894 | 9.38497066 | 4.00000000 |
| 0.783399999 | 1.63223302 | 4.25715160 | 4.00000000 |
| 0.940599978 | 6.63401318 | 17.7794666 | 4.00000000 |
| 1.29740000 | 5.49634790 | 15.2450886 | 4.00000000 |
| 1.71099997 | 1.81207728 | 4.86538553 | 4.00000000 |
| # Positive Lines | : | | |
| 0.470999986 | 1.57258713 | 4.53434849 | 4.00000000 |
| 0.685000002 | 1.54889691 | 4.23500299 | 4.00000000 |
| | | | |

Tentative Redshift IDs

| Possible Redshift: z = | = 0.611312 | 509 | | | |
|--------------------------------------|--|---------|---------------|-------------|----------|
| ==================================== | ====================================== | 0.573 | 949516 · lon· | 07 | |
| 2 : Eobs = 0.4 | 13399994 | ; E0 = | 0.666116595 | ; lon: 07 | |
| 3 : Eobs = 0.5 | 73800027 | ; E0 = | 0.924571157 | ; Ion: Ne9 | |
| 4 : Eobs = 0.6 | 23799980 | ; E0 = | 1.00513673 | ; Ion: Fe21 | |
| Possible Redshift: z : | = 0.388363 | 600 | | | |
| 1: Eobs = 0.413399 | 994 ; E0 = | 0.573 | 949516 ; lon: | 07 | |
| 2: Eobs = 0.5 | 03000021 | ; E0 = | 0.698346913 | ; Ion: O7 | |
| 3: Eobs = 0.6 | 23799980 | ; E0 = | 0.866061211 | ; Ion: Fe18 | |
| 6: Eobs = 0.6 | 41399980 | ; E0 = | 0.890496373 | ; Ion: Fe18 | |
| | | | | | |
| Possible Redshift: z : | = 0.263649 | 344 | | | |
| 1: Eobs = 0.454200 | 000 ; E0 = | 0.573 | 949516 ; lon: | 07 | |
| 2: Eobs = 0.5 | 64999998 | ; E0 = | 0.713961899 | ; Ion: 07 | |
| 3 : Eobs = 0.7 | 10600019 | ; E0 = | 0.897949219 | ; Ion: Fe17 | |
| 4 : EODS = 0.7 | | ; EU = | 0.898960114 | ; ION: Fe17 | <u> </u> |
| Possible Redshift: 7 : | = 0.141052 | 723 | | | 1 |
| | ======= | | | | |
| 1: Eobs = 0.503000 |)21 ; E0 = | 0.573 | 949516 ; lon: | 07 | |
| 2: Eobs = 0.5 | 73800027 | ; E0 = | 0.654736102 | ; Ion: 08 | |
| 4:EODS = 0.0 5:Fobs = 0.7 | 83399990 | ; EO = | 0.711788654 | ; Ion: 07 | |
| 6: Eobs = 0.9 | 40599978 | ; E0 = | 1.07327414 | ; Ion: Ne9 | |
| | | | | | |
| Possible Redshift: z : | = 1.583981 | .51E-02 | | | |
| 1: Eobs = 0.564999 | 998 : E0 = | 0.573 | 949516 ; Ion: | 07 | |
| 2 : Eobs = 0.6 | 40600026 | ; E0 = | 0.650747001 | ; Ion: O8 | |
| 4 : Eobs = 0.6 | 41399980 | ; E0 = | 0.651559651 | ; Ion: 08 | |
| 6: Eobs = 0.7 | 10600019 | ; E0 = | 0.721855819 | ; lon: 07 | |
| 7 : EODS = 0.7 | | ; E0 = | 0.722668409 | ; Ion: 07 | |
| Possible Redshift: z : | = 2.604722 | 98E-04 | | | |
| | | | | | |
| 1: Eobs = 0.573800 |)27 ; E0 = | 0.573 | 949516 ; lon: | 07 | |
| 2:EODS = 0.7 3:Eobs = 0.7 | 10600019 | ; E0 = | 0.710785091 | ; Ion: 07 | |
| | | , LO – | | | |
| Possible Pedshift: 7 | 5 ///80 | 0265-07 |) | | |
| | | 020E-07 | - | | |
| 1: $Eobs = 0.6069999$ | 993 ; E0 = | 0.573 | 949516 ; lon: | 07 | |
| 2 : Eobs = 0.9 | 40599978 | ; E0 = | 0.889385343 | ; lon: Fe18 | |
| Possible Redshift: z : | = -7.99141 | 526E-02 | 2 | | |
| 1: Fobs - 0.622700 | | 0 5 7 2 | 040516 · lon: | 07 | |
| 2: Eobs = 0.023799 | 10600019 | : E0 = | 0.653813004 | : lon: 08 | |
| 1 - Eobs - 0.7 | 11200072 | , 50 | 0.0535015000 | , 1 | |

6 : Eobs = 0.783399999 ; E0 = 0.720795274 ; Ion: O7 7 : Eobs = 0.940599978 ; E0 = 0.865432739 ; Ion: Fe18

BLS Example: Line and Redshit IDs



z=0.141

0.4

0.5

0.6

Energy (in keV)

z=0.611

0.8

0.9 1

0.7

-15

-20 ⊾ 0.3

| Possible Redshift: z = 0.611312 | 509 | | |
|---------------------------------|--------|---------------|-------------|
| | | | |
| 1: Eobs = 0.356200010 ; E0 = | 0.5739 | 949516 ; lon: | 07 |
| 2 : Eobs = 0.413399994 | ; E0 = | 0.666116595 | ; lon: 07 |
| 3 : Eobs = 0.573800027 | ; E0 = | 0.924571157 | ; Ion: Ne9 |
| 4 : Eobs = 0.623799980 | ; E0 = | 1.00513673 | ; lon: Fe21 |

| Possible Redshift: z = 0.141052723 |
|---|
| ======================================= |
| 1: Eobs = 0.503000021 ; E0 = 0.573949516 ; Ion: O7 |
| 2 : Eobs = 0.573800027 ; E0 = 0.654736102 ; Ion: O8 |
| 4 : Eobs = 0.623799980 ; E0 = 0.711788654 ; Ion: O7 |
| 5 : Eobs = 0.783399999 ; E0 = 0.893900692 ; Ion: Fe17 |
| 6 : Eobs = 0.940599978 ; E0 = 1.07327414 ; Ion: Ne9 |

BLS: TO dO (depending on funding, personnel, etc.)

- Update/Integrate (forbidden/intercombination, innershell resonant ground-state and metastable transitions) Atomic Database
- Improve Line/Redshift ID routine: needs to be smarter (AI training might help): e.g. starting from strongest and looking for right strength ratios (including saturation)
- Python interface to directly call BLS from within Sherpa | XSPEC
- GUI interface to locally iterate the procedure and plot results

14.8: Time-Evolving Photo-Ionization Device (TEPID: Luminari+22)

- Solves the system of 1st-order differential equations of time-evolving ionization balances following ionizing intensity variations with time
- Uses adaptive time-resolution algorithm to speed-up calculation
- Currently considers only photo-ionization and radiative recombinations and approximates on heating-cooling balance
- Includes radiative-transfer
- Applies to all variable ionizing sources: e.g. transients (GRBs) and AGNs
- Interfaces with Phase (Krongold+03), in XSPEC|Sherpa to produce mock spectra and fit data

$$\frac{dn_{Xi}}{dt} = -[F_{Xi} + \alpha_{\rm rec}(X^{i-1}, T_e)n_e]n_{Xi} + F_{Xi-1}n_{Xi-1} + \alpha_{\rm rec}(X^i, T_e)n_e n_{Xi+1}.$$
 (4)

$$t_{eq} \approx \left\{ (\alpha_{rr}(X^i, T_e)n_e) \cdot \left(\frac{\alpha_{rr}(X^{i-1}, T_e)}{\alpha_{rr}(X^i, T_e)} + \frac{n_{X^{i+1}}}{n_{X^i}} \right) \right\}^{-1}$$

TEPID: Adaptive-Resolution Algorithm



Up-and-Down case



TEPID: GRB-case: Stratified-Ionization



TEPID: GRB-case: Spectra vs Local Density



TEPID: AGN-case: Ionization Balance vs Time



Fig. 9: AGN lightcurve



TEPID: AGN-case: Time-Resolved Spectra





TEPID: To do

- Include time-evolving population-level calculation and metastable transitions (to further constrain density)
- Include proper heating/cooling calculation
- Include additional ionization/recombination mechanisms (i.e. transition rates)
- Improve atomic-database
- Include proper pre-burst star-forming region ionization conditions based on simulations (Luminari, Graziani, in prep.)